

# pLISA: a parallel Library for Identification and Study of Astroparticles and its application to KM3NeT

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for C.B. \*, C. De Sio \*\*, R. Coniglione \*\*\**

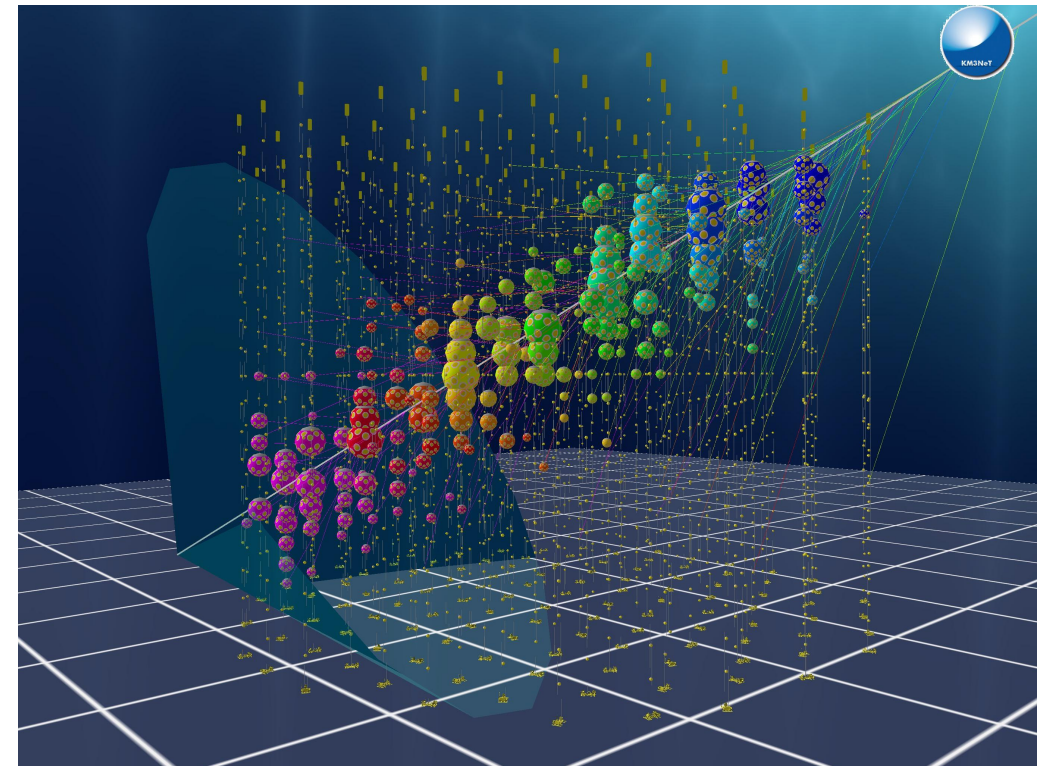
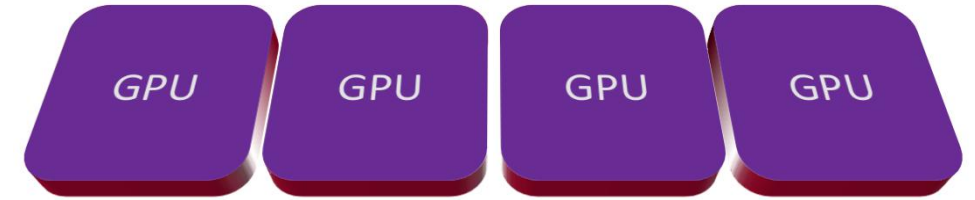
*The new Era of Multimessenger Astrophysics*

*Groningen, March 28th 2019*

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- \*\* Formerly at University of Salerno and INFN “Gruppo collegato di Salerno”
- \*\*\* INFN Laboratori Nazionali del Sud - Catania

# Motivations for pLISA

- Leverage parallel computing based on GPUs
- Primary particle reconstruction for event-based detectors
- Set up “continuously training farms”



# pLISA structure

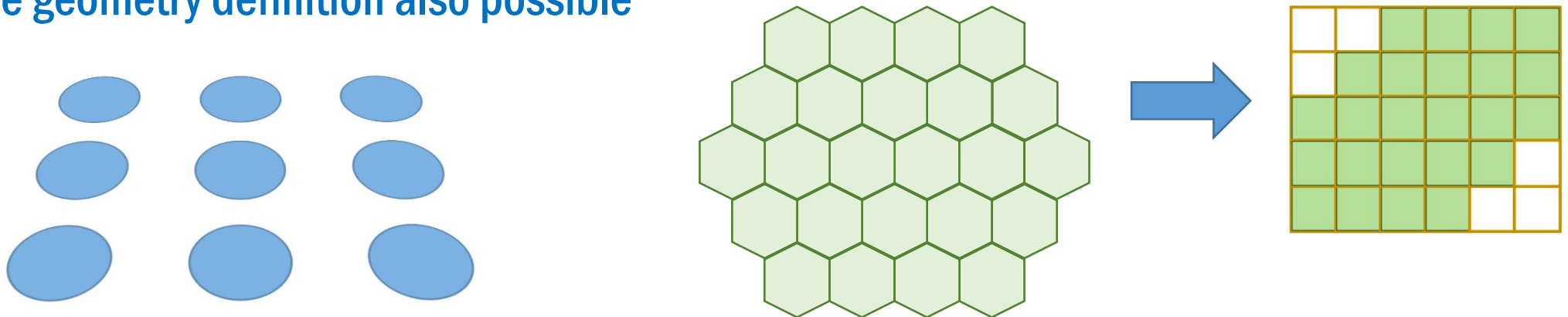
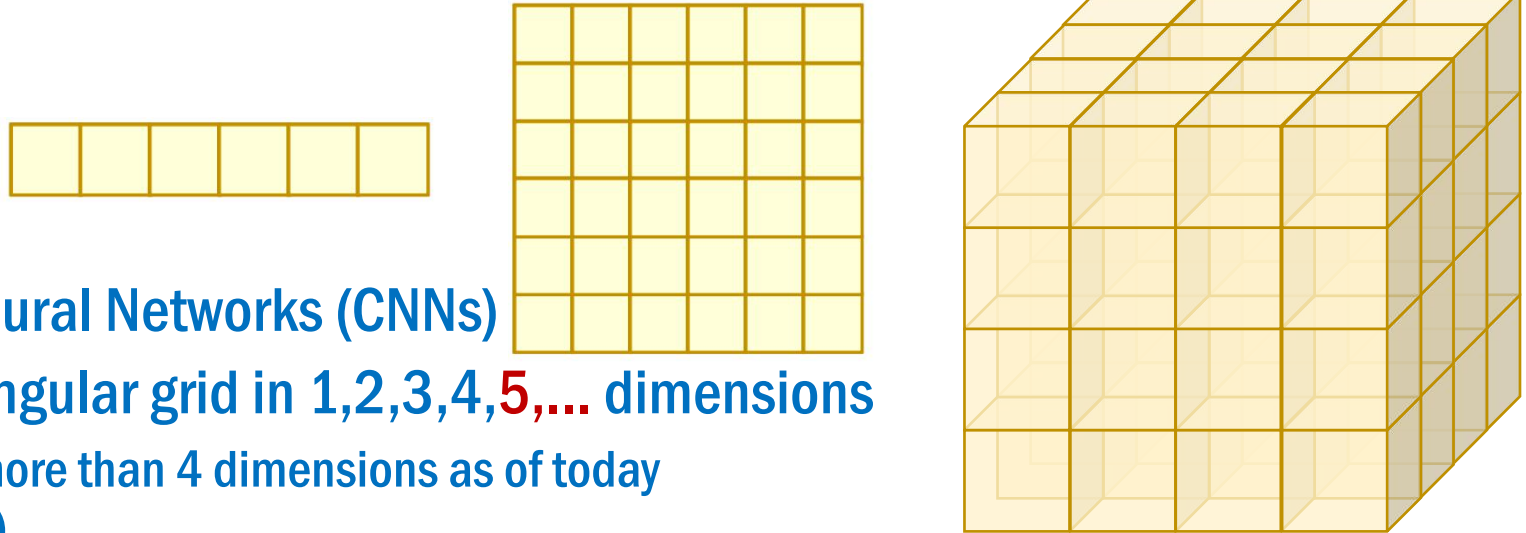
- Set of Python scripts (conveniently used in Jupyter notebooks)
- Scikit Learn  
(data managed by image handling techniques)
- TensorFlow  
(differentiable programming and machine learning)
- Keras (high-level neural network software)



# pLISA concepts

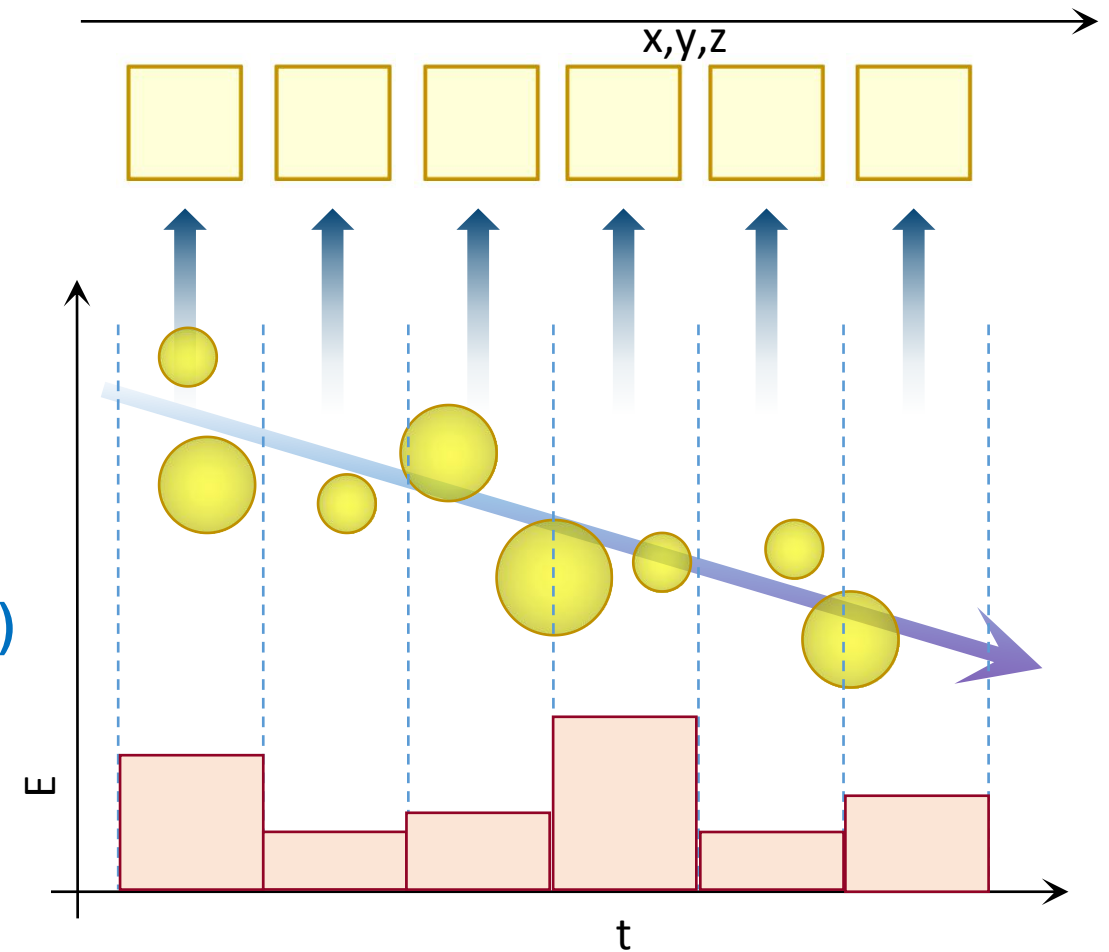
## • Detector representation

- Suitable for Convolutional Neural Networks (CNNs)
- Must be mappable to a rectangular grid in 1,2,3,4,5,... dimensions
  - CNN stack does not support more than 4 dimensions as of today (hope to increase in the future)
- Piecewise geometry definition also possible



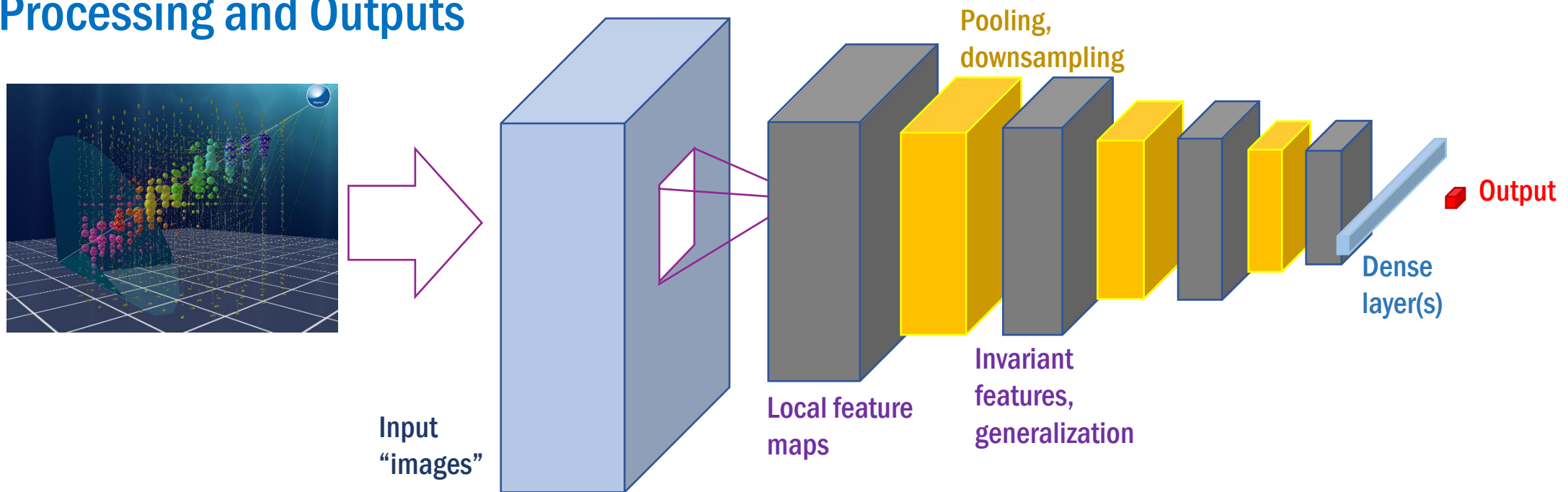
# pLISA concepts

- Event representation
  - Spatial discretisation (implied in discretised detector structure)
  - Time discretisation (“hits” in sensing elements are binned in time, by count or deposited energy)



# pLISA concepts

- Processing and Outputs



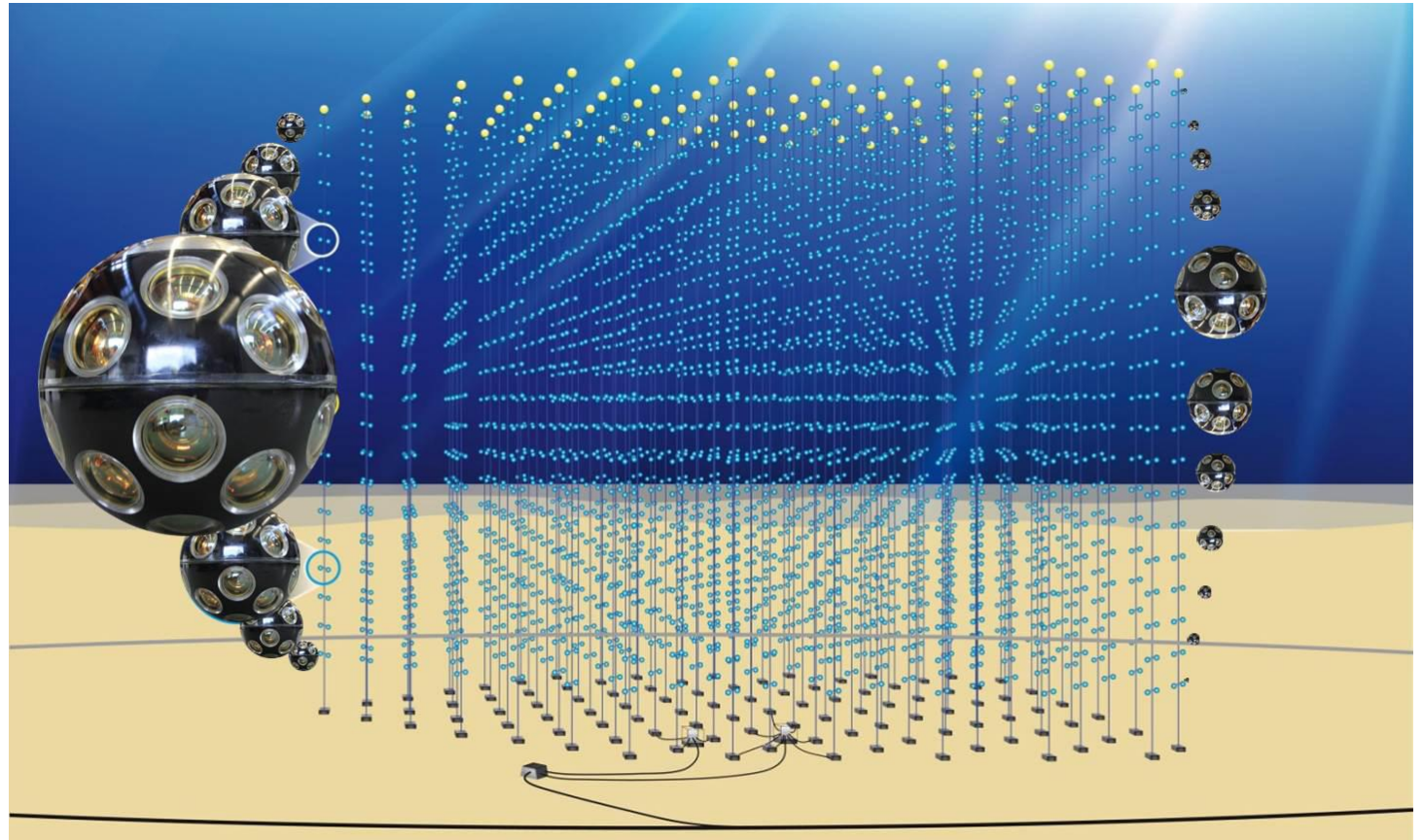
**No prerequisite reconstruction output - depends on no other software**

CNN could instead help/support standard reconstruction algorithms with first guess



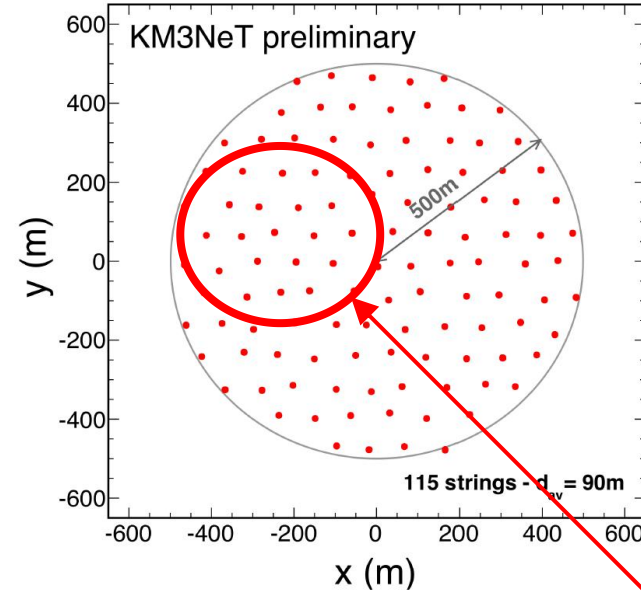
# Application to KM3NeT

- KM3NeT Water Cherenkov in deep sea
- Building block: 115 DUs
- Detection Unit (DU):  
18 evenly spaced DOMs
- Digital Optical Module (DOM):  
31 PMTs (+ other instruments)

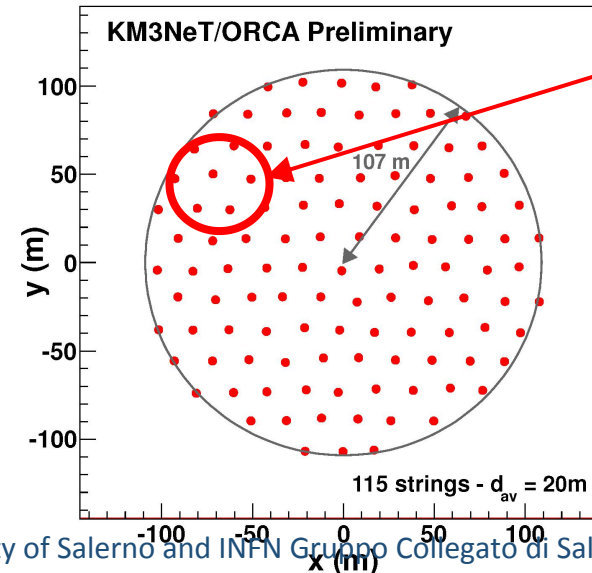
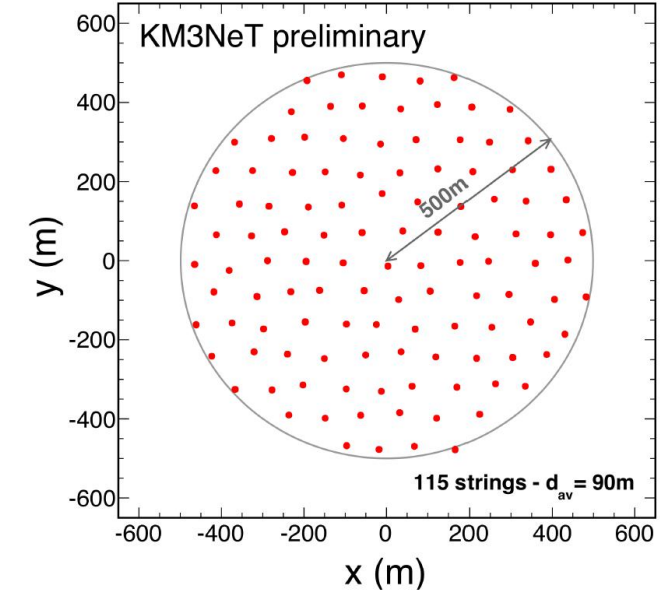


# Application to KM3NeT

- **Astroparticle Research with Cosmics in the Abyss**  
km<sup>3</sup> size building blocks (2× )  
z DOM spacing 36 m
- **Oscillation Research with Cosmics in the Abyss**  
1 building block  
z DOM spacing 9 m



+



Phase1 (fully funded - deploy 2019)

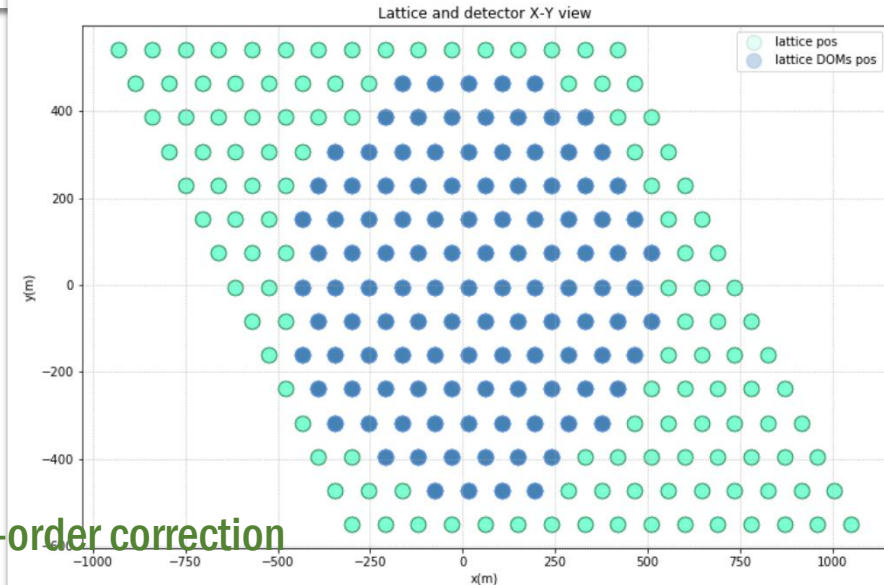
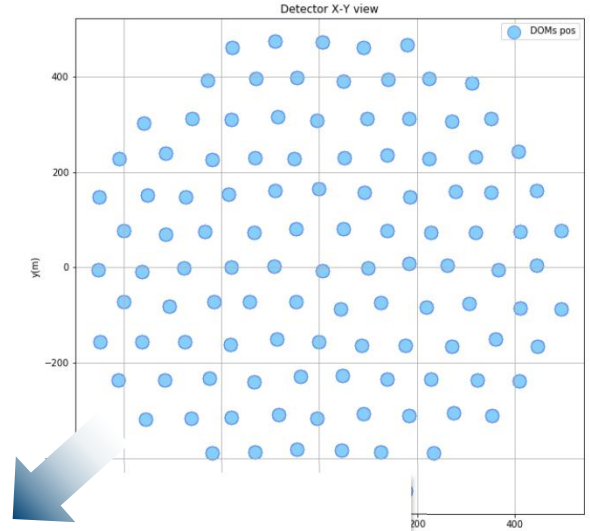
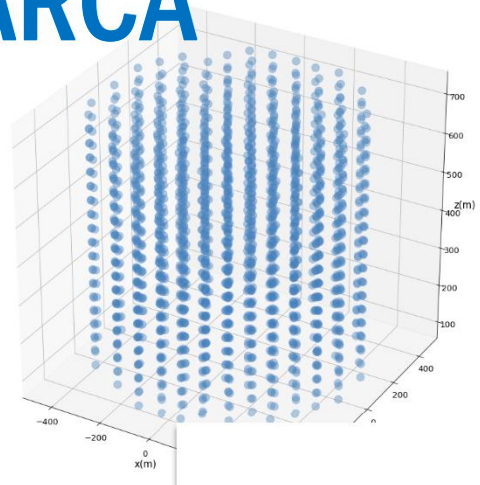
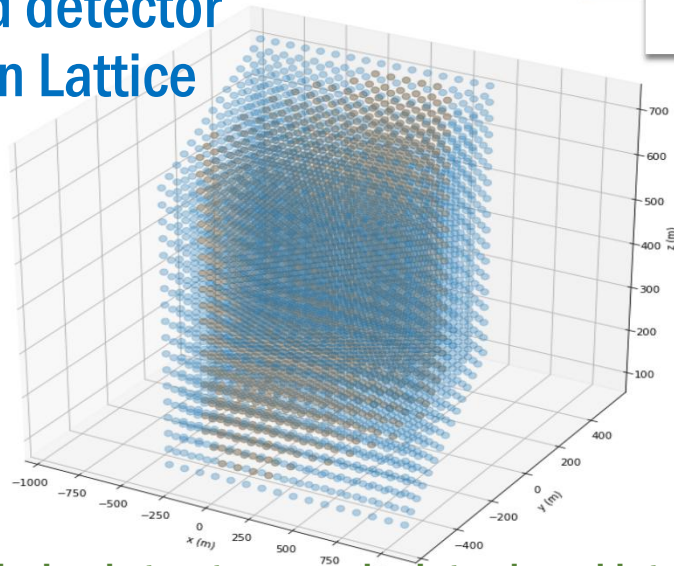
Phase 2 (partially funded - deploy 2019/21)

KM3NeT 2.0 Letter of Intent:  
arXiv:1601.07459 and  
J.Phys. G43 (2016) 084001



# Application to KM3NeT - ARCA

- XYZ detector regularisation
  - exactly 90m spaced in (X,Y)
  - exactly 36m spaced in Z
  - Regularised detector contained in Lattice

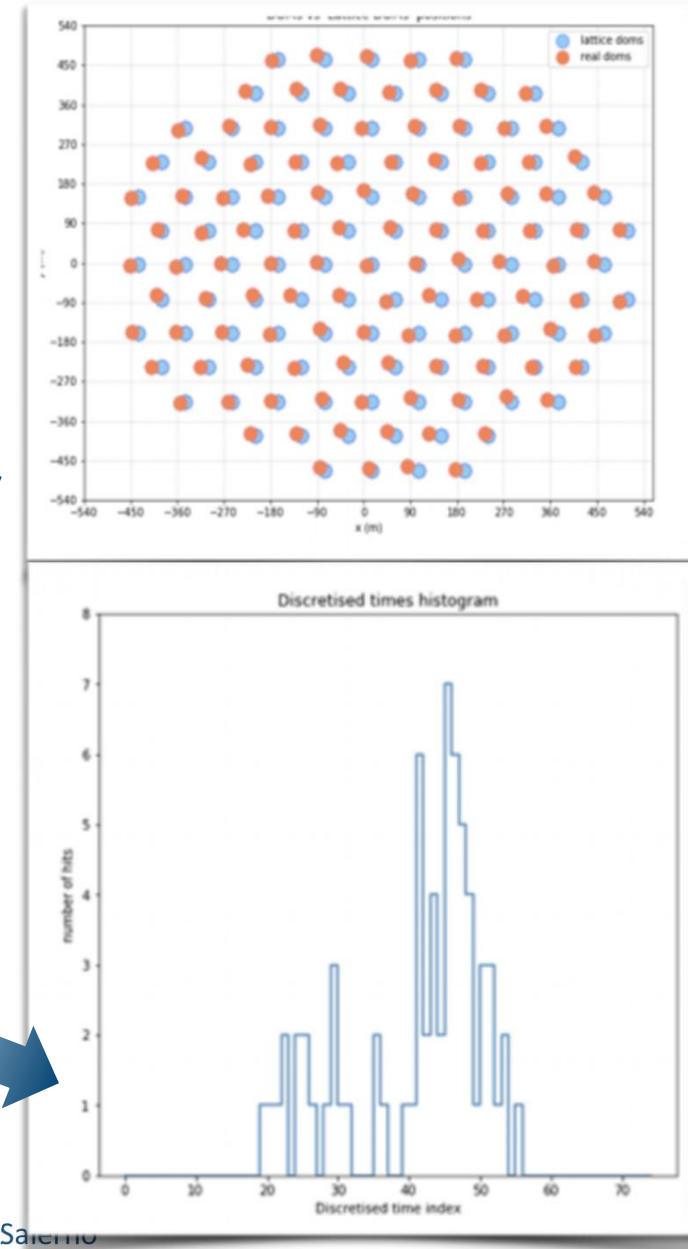


Deviation from regularised structure can be introduced later as a next-order correction

# Application to KM3NeT - ARCA

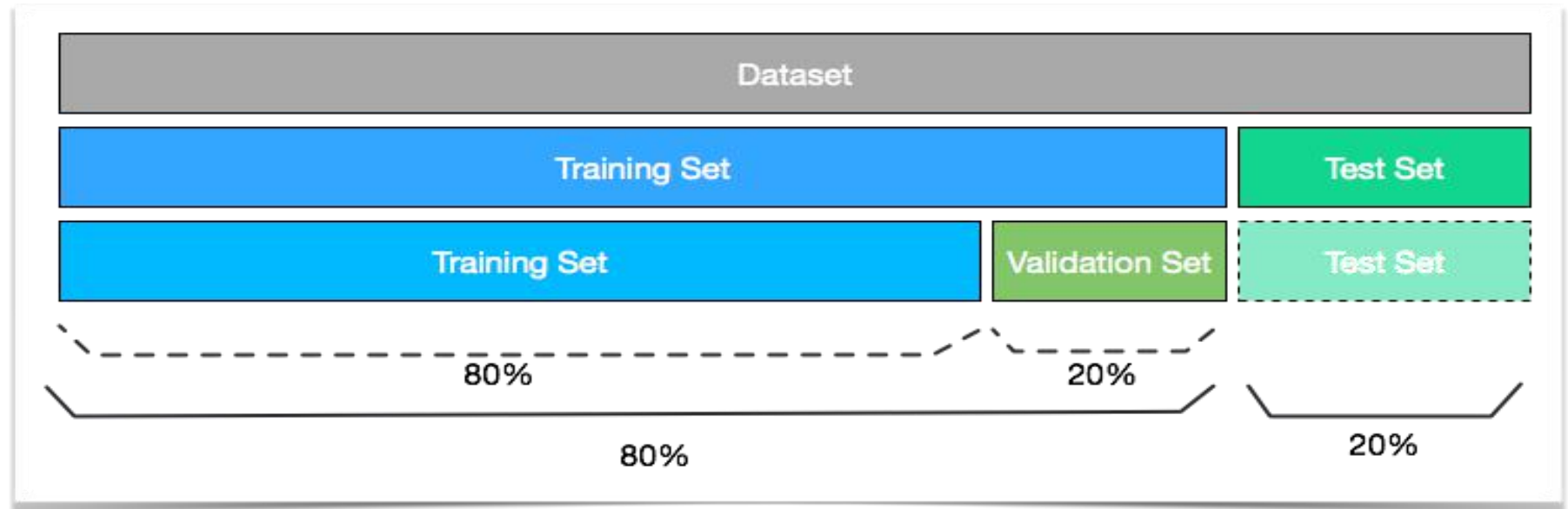
- Event representation

- Reduce data sparsity: ignore PMT orientation and consider only DOM
- Force DOM positions onto regularised lattice
- DOM mapped to a position in  $16 \times 15 \times 18$  lattice
- Hit time binned to 12 ns bin  $\Rightarrow$  75 bins / event
- Matrix structure of any event (TXYZ):  $[75 \times 16 \times 15 \times 18]$   
4D image (or 3D movie)



# Application to KM3NeT - ARCA

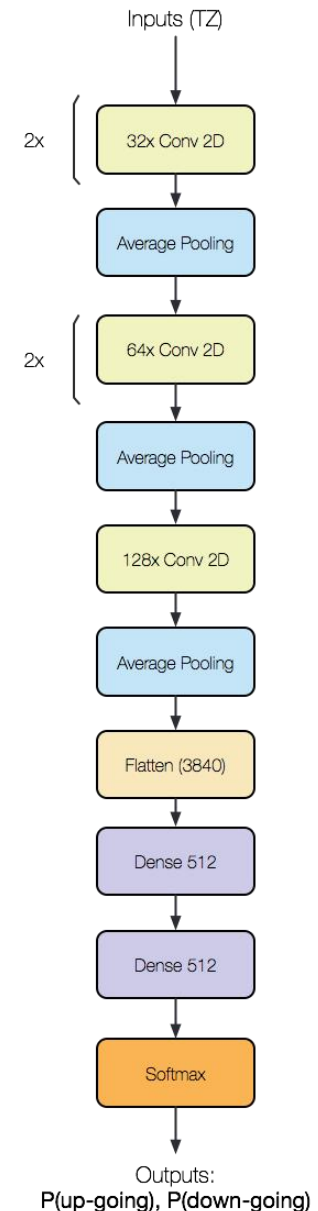
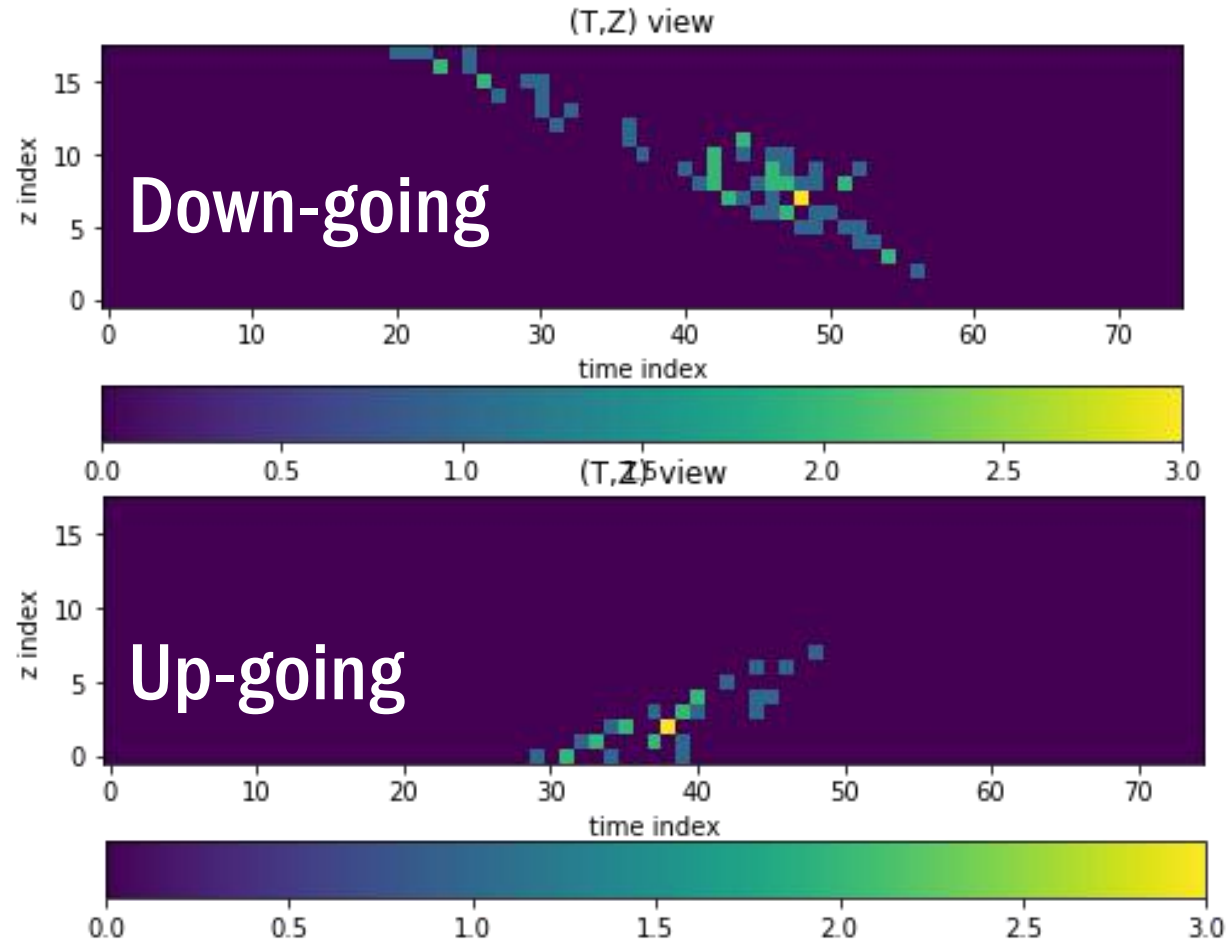
- 258,879 events
- $\nu_e$ CC and  $\nu_\mu$ CC



- Training no longer than a few hours
- Slow models also tested (training for several days) but discarded (no real improvement)

# KM3NeT - Up/Down-going classification

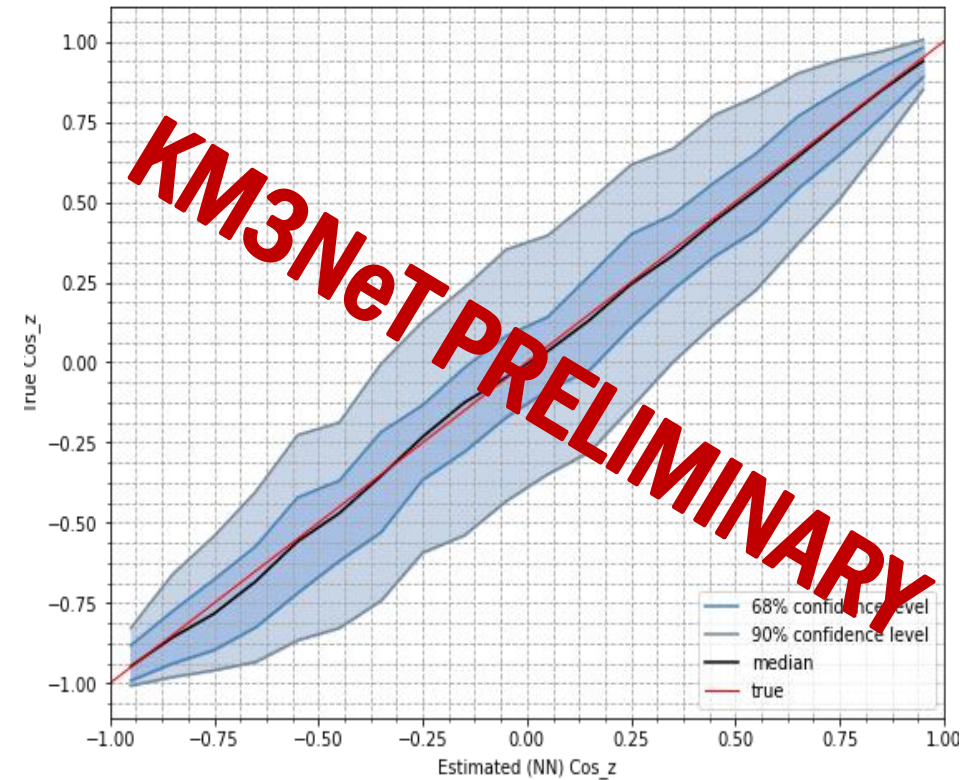
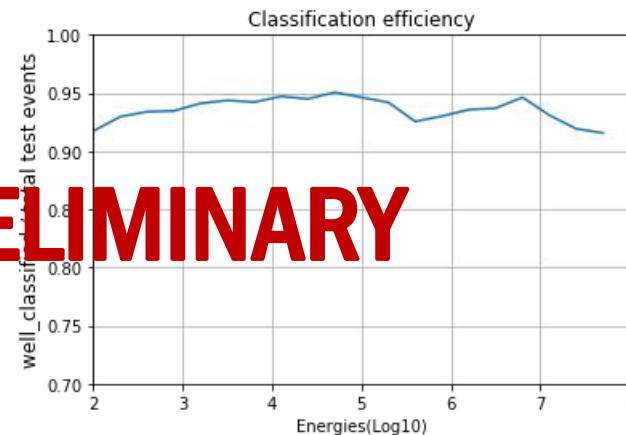
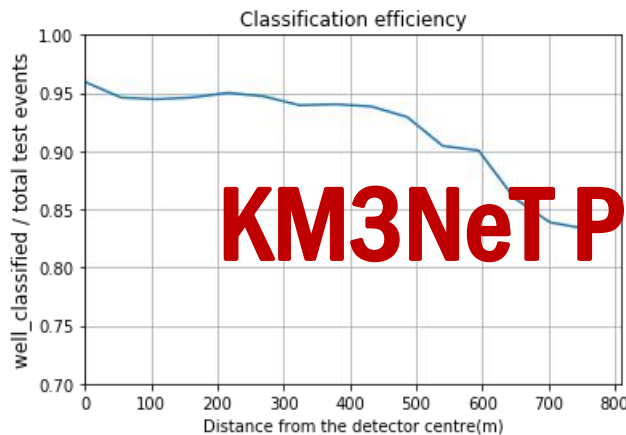
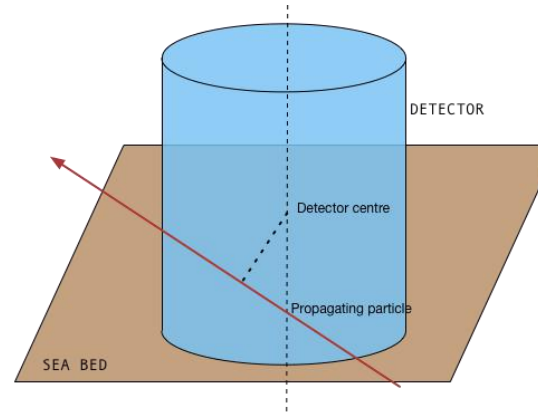
- Colour code:  
hits/DOM/time bin
- XY found to be irrelevant for  
this task





# KM3NeT - Direction(Z) & Up/Down-going classification

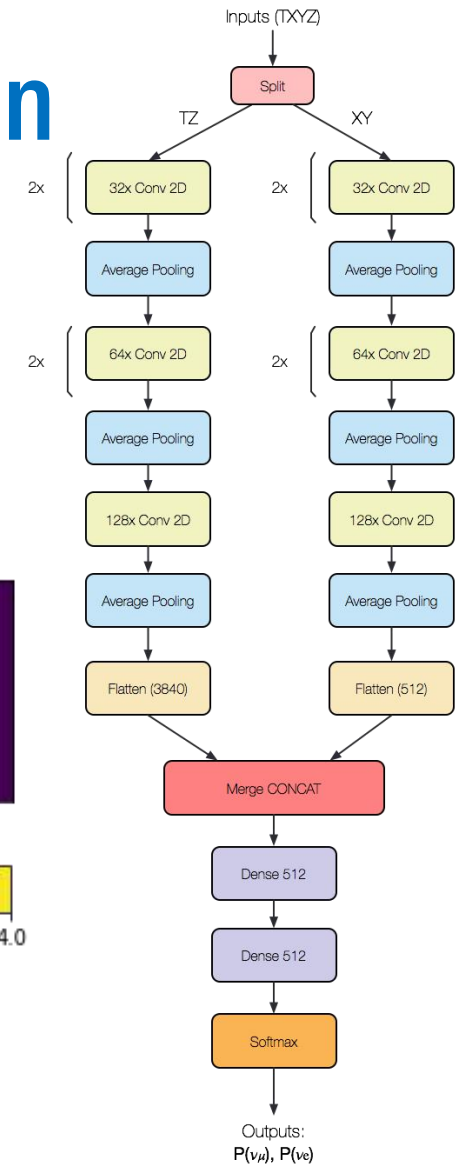
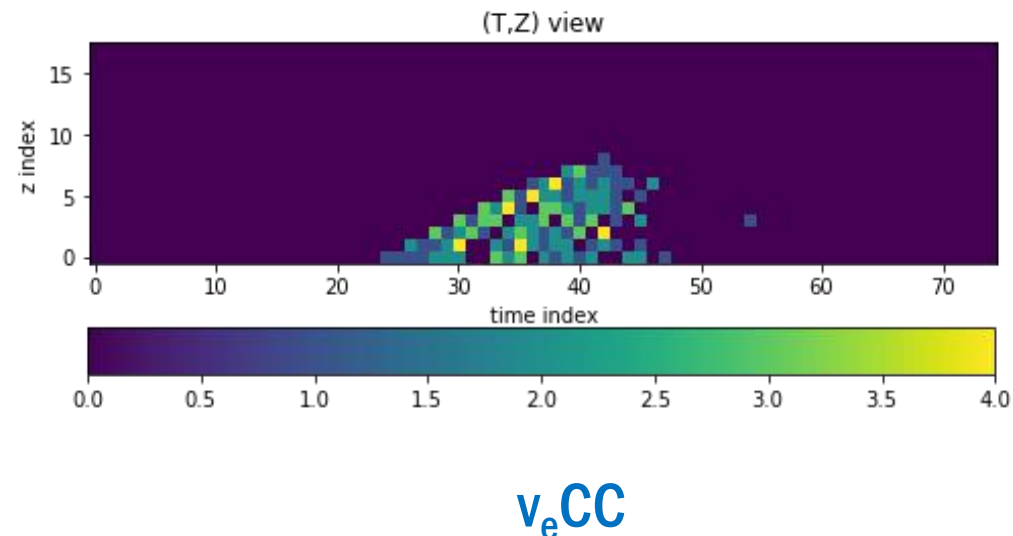
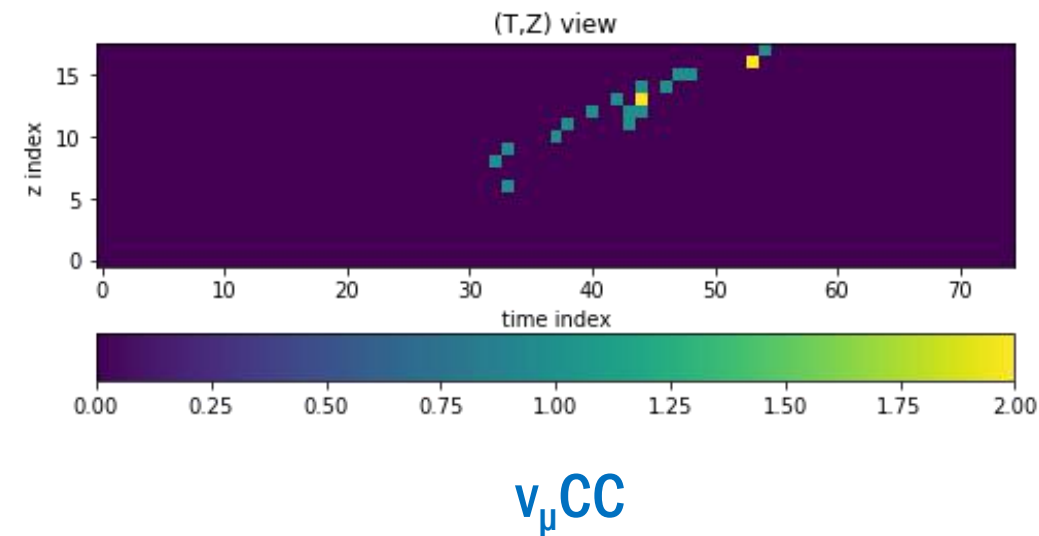
- $\cos \theta_z$  estimated by network
- Performance depends on track length in detector



$RMS(\cos \theta_{z,est} - \cos \theta_{z,true}) = 0.001$  for traditional algorithms but using quality cuts, 0.002 for CNN - but no specific optimization done, and works on all events

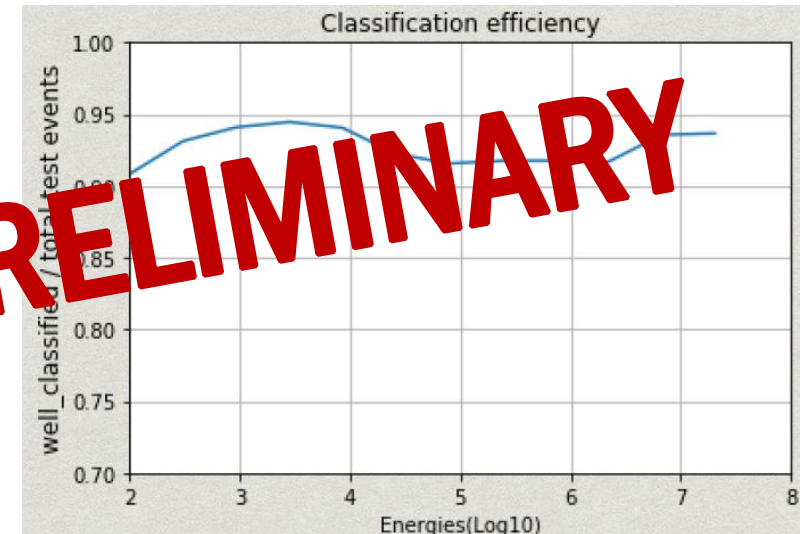
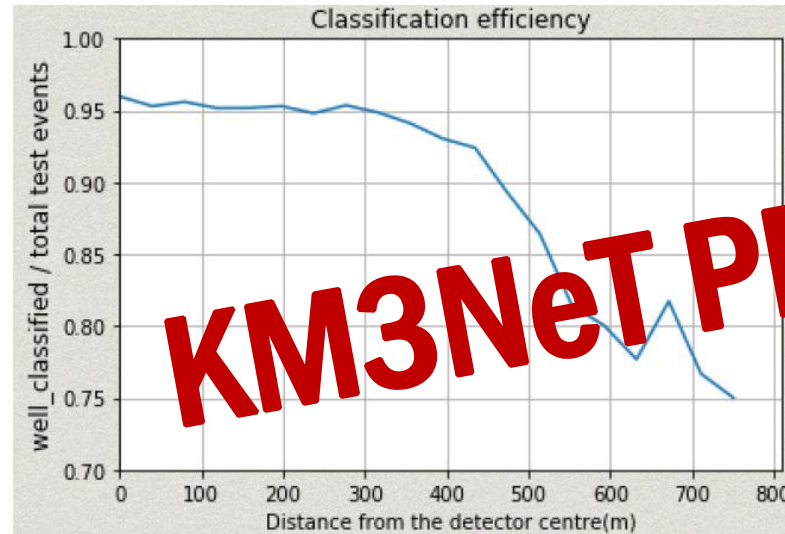
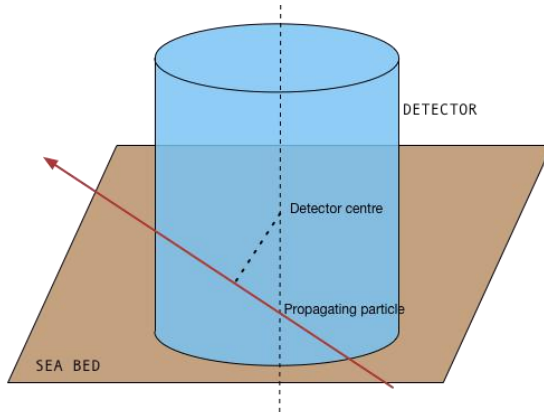
# KM3NeT - Event type ( $\nu_e$ CC/ $\nu_\mu$ CC) classification

- XY shape and TZ evolution analysed separately in the first stages, merged at later stage after extraction of features



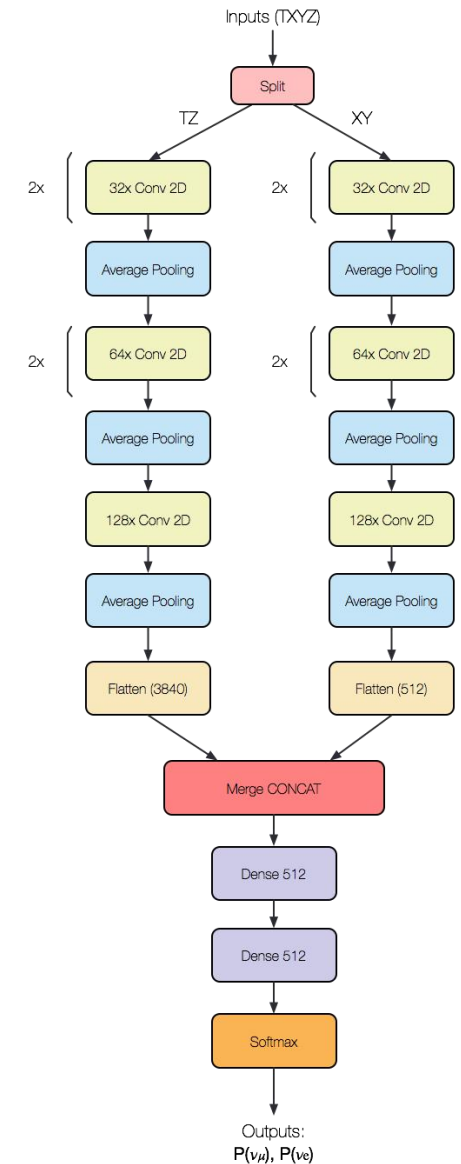
# KM3NeT - Event type ( $\nu_e$ CC/ $\nu_\mu$ CC) classification

- Classification efficiency found to depend on track length in detector
- Energy dependency milder



# KM3NeT - $\nu$ Energy estimation

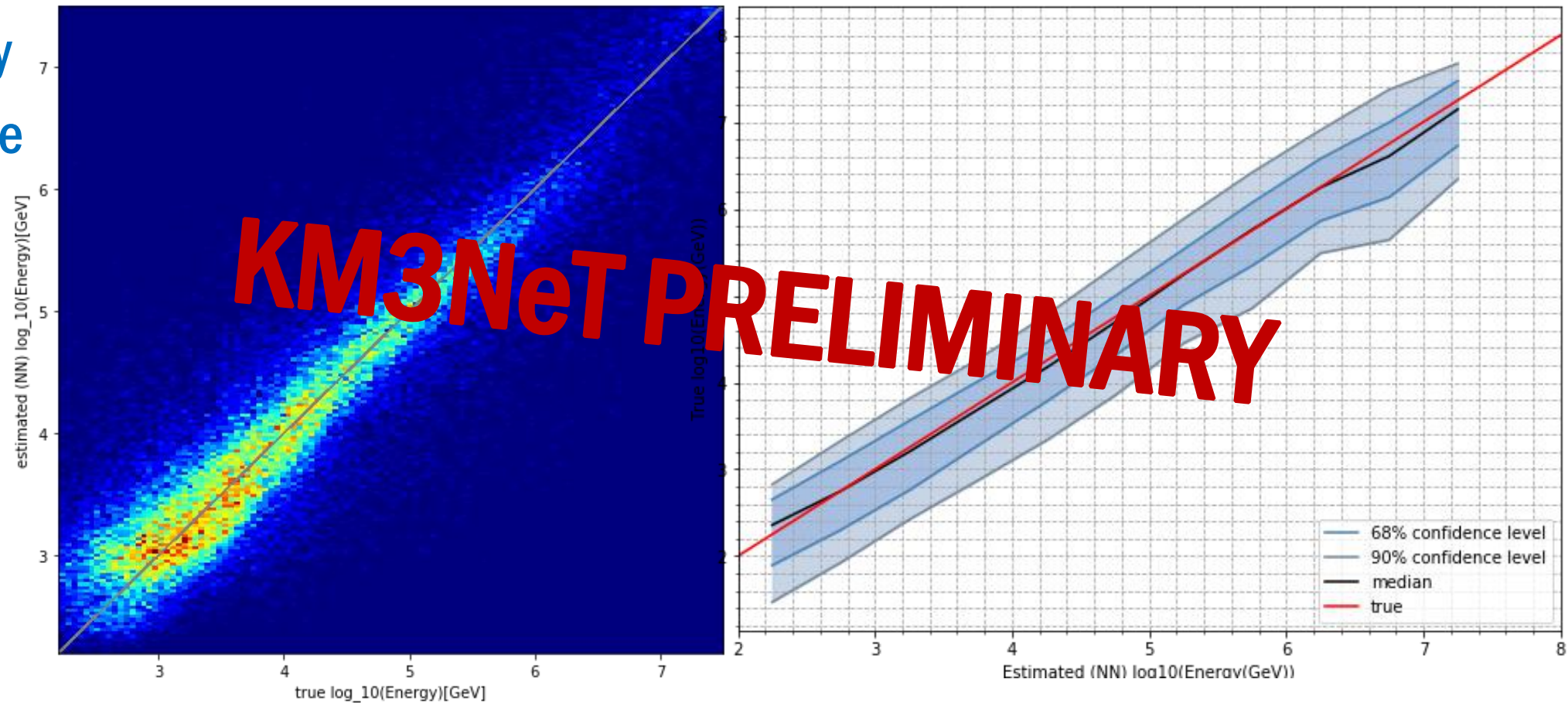
- Events are not pre-classified with respect to particle ID
- The CNN does not have any other information than hits
- $\nu_e$ CC and  $\nu_\mu$ CC handled together and mixed in training/test/validation samples





# KM3NeT - $\nu$ Energy estimation

- Good linearity
- Could improve with more statistics at high energy



# Conclusions

- pLISA is a convenient framework to flexibly develop neural networks for astroparticle identification and study
- Based on widespread technologies
- Usable output can already be produced
- Can already provide a cross-check for “traditional” reconstruction algorithms



<https://baltig.infn.it/bozza/plisa/>

# Outlook

- Improve and clean up library API
- Provide support for problems with higher dimensionality
- Application side: try to include all neglected ingredients for real situations, including efficiency, irregular shapes, noise, detector distortion, etc.

